

Nuclear Astrophysics Opportunities with the Modular Total Absorption Spectrometer (MTAS)

B.C. Rasco

Workshop on Nuclear Astrophysics
Opportunities at ATLAS 2019

July XX, 2019

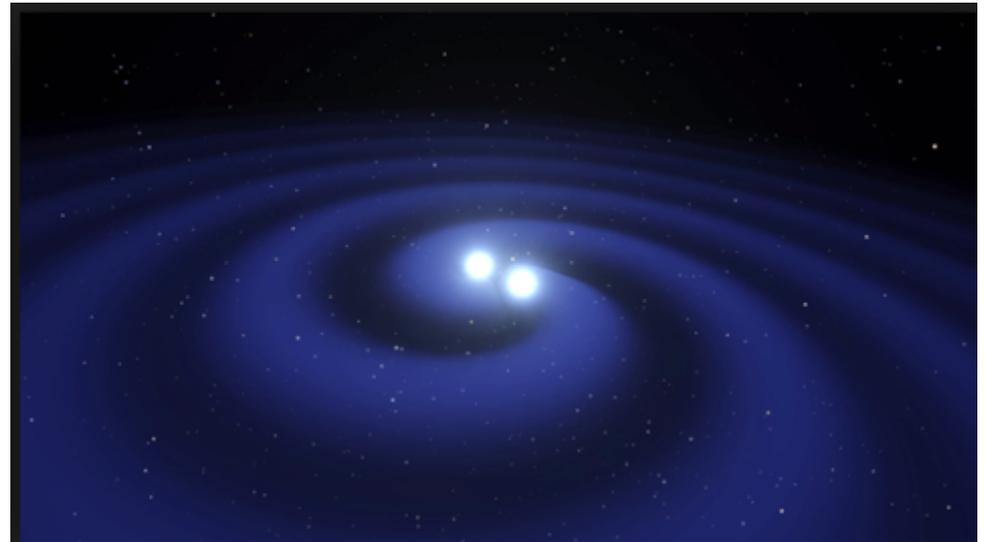
Nuclear Astrophysics?

What qualifies as nuclear astrophysics?

What can the Modular Total Absorption Spectrometer (MTAS) measure?



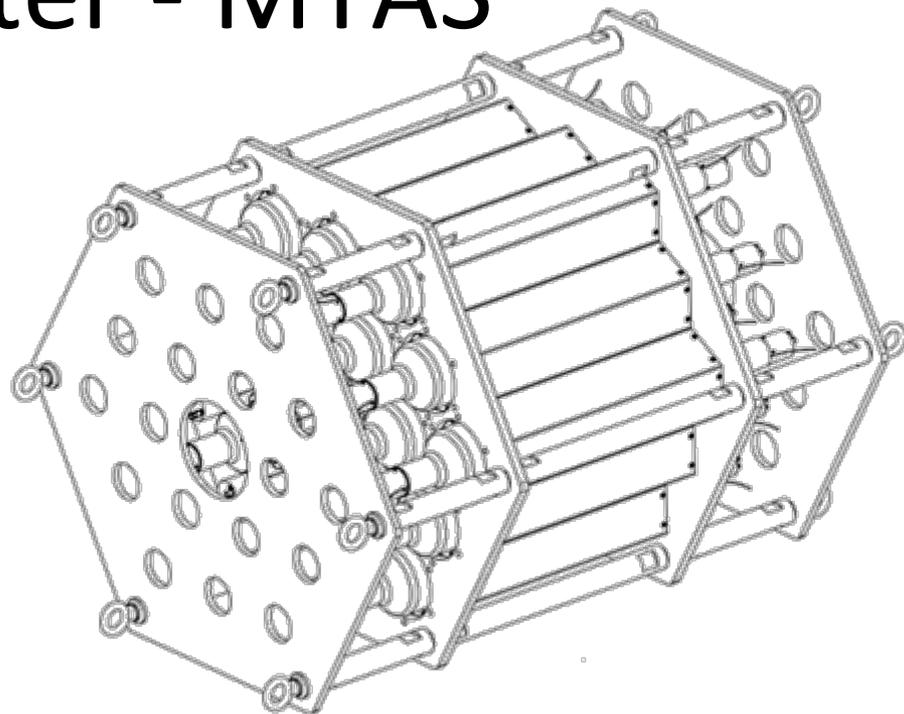
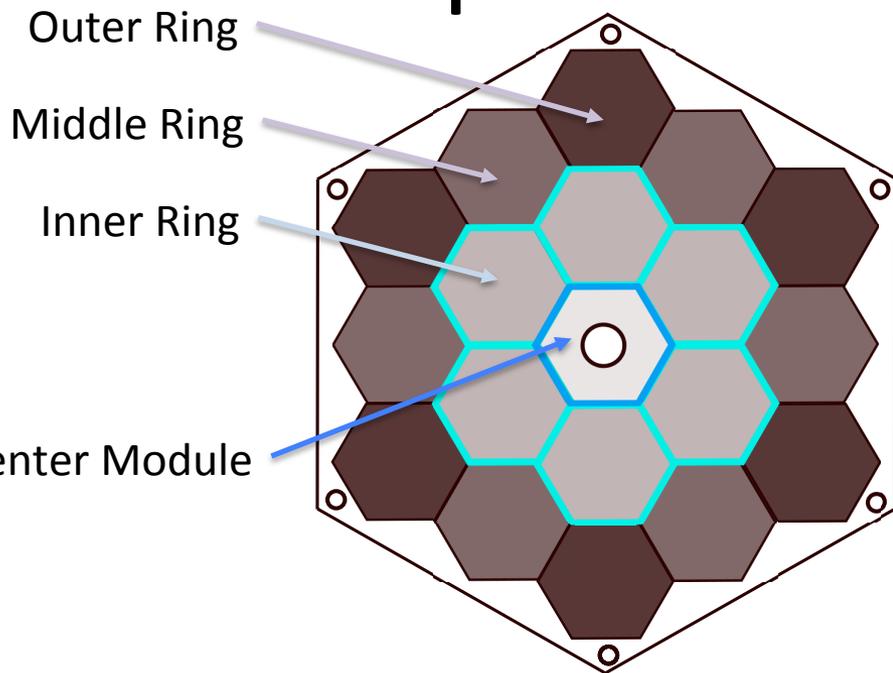
Dark Matter



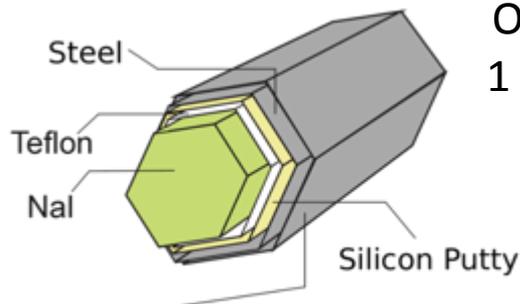
r Process - Explosion
r Process - Freeze Out

ESA

The Modular Total Absorption Spectrometer - MTAS



MTAS: 18 - 8" x 7" x 21" (20cm x 17.8cm x 53.3cm) hexagon NaI(Tl) modules
 Organized in 3 Rings of 6 modules each (Inner, Middle, and Outer)
 1 - Center module, same dimensions but with a 2.5" diameter hole
 Over 1 ton of NaI(Tl)!



Over 5 tons of lead shielding + neutron shielding

Other total absorption spectrometers include the TAS at ISOLDE, Lucrecia, TAS at GSI (now at UML), SuN (MSU), DTAS (Valencia, Jyvaskyla), HECTOR(ND).

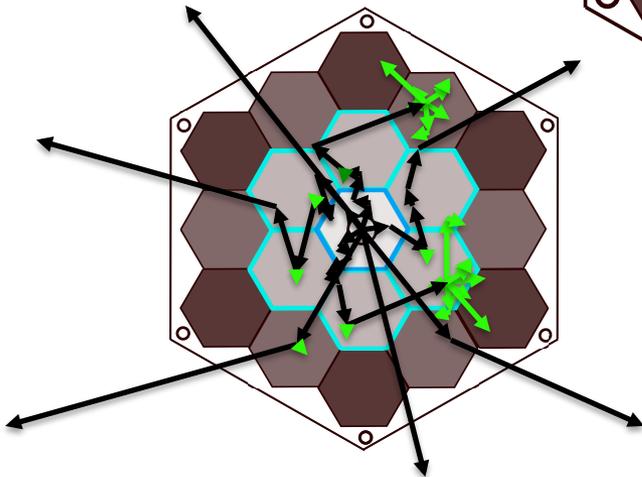
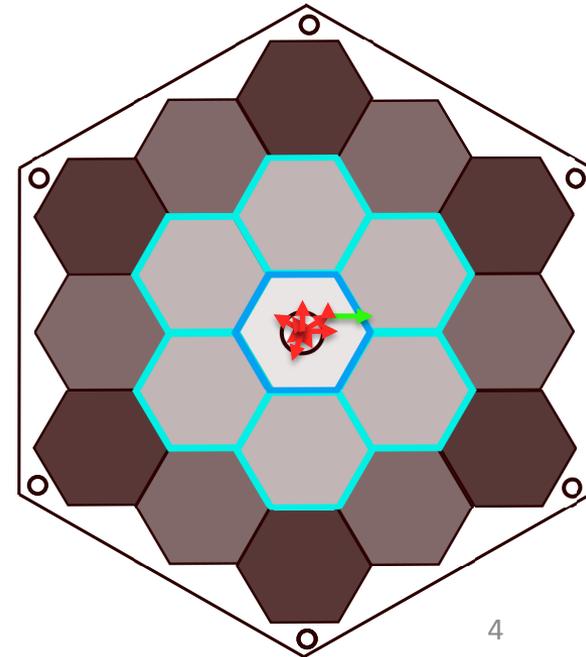
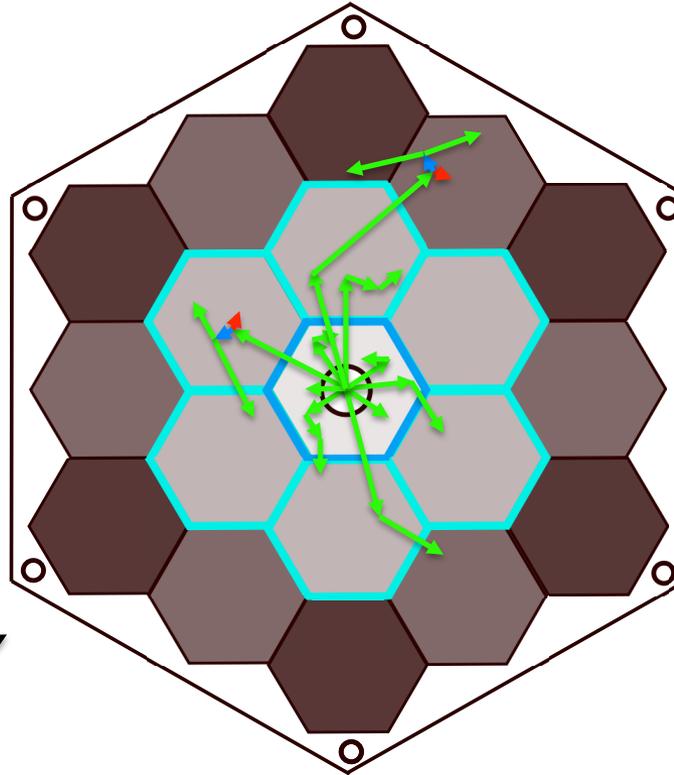
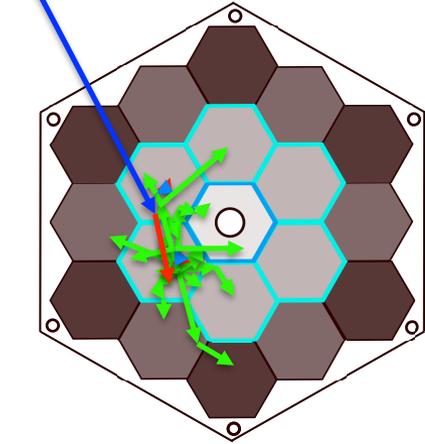
What MTAS Detects

γ s in MTAS

β s in MTAS

μ s in MTAS

Neutrons in MTAS



What MTAS Measures

For neutron rich nuclei, β s, γ s, and neutrons will be present in almost every decay. MTAS can measure them all. By measuring all as a function of energy, we can measure the β -strength function in one experiment.

MTAS γ s

Detects single γ s with a $\sim 98\%$ efficiency

Detects **level** feeding, which is directly relatable to Gamow-Teller strength

MTAS β s

Detects β s with $E > 2$ MeV

MTAS Neutrons

Detects neutrons with $\sim 15\%$ efficiency

Measures neutron energy ~ 250 keV FWHM

Detects very low energy neutrons

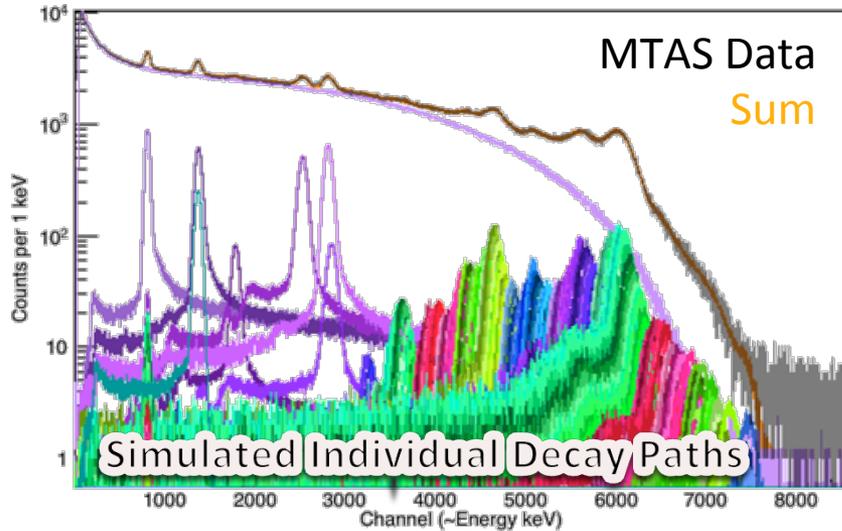
Detects neutron fine structure (γ s following β -delayed neutron decay) feeding intensities

Measuring all three with one detector in one experiment is efficient.

MTAS is a unique and powerful device.

Measuring β s: ^{92}Rb

$Q_\beta = 8095(6)$ keV $T_{1/2} = 4.48(3)$ s



B.C. Rasco, *et al.*, PRL **117**, 092501 (2016)

Ground state feeding insensitive to exact decay pattern from higher states

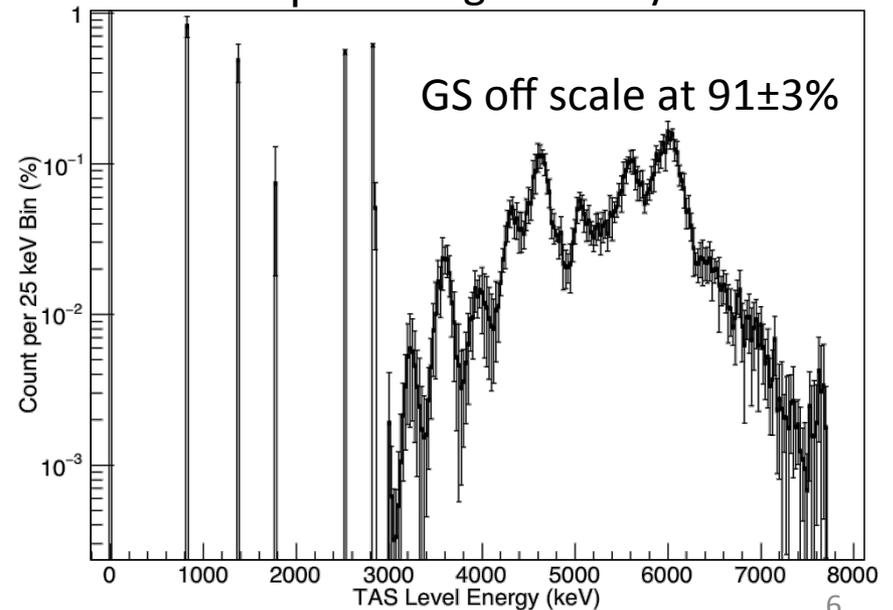
ENSDF Ground State Feeding: $95.2 \pm 0.7\%$
(up from $50 \pm 18\%$ in 2007)
and $87.5 \pm 2.5\%*$

*A.-A. Zakari-Issoufou *et al.*, PRL **115**, 102503

MTAS Ground State Feeding: **$91 \pm 3\%$**

Our uncertainty mainly from ground state β simulation.

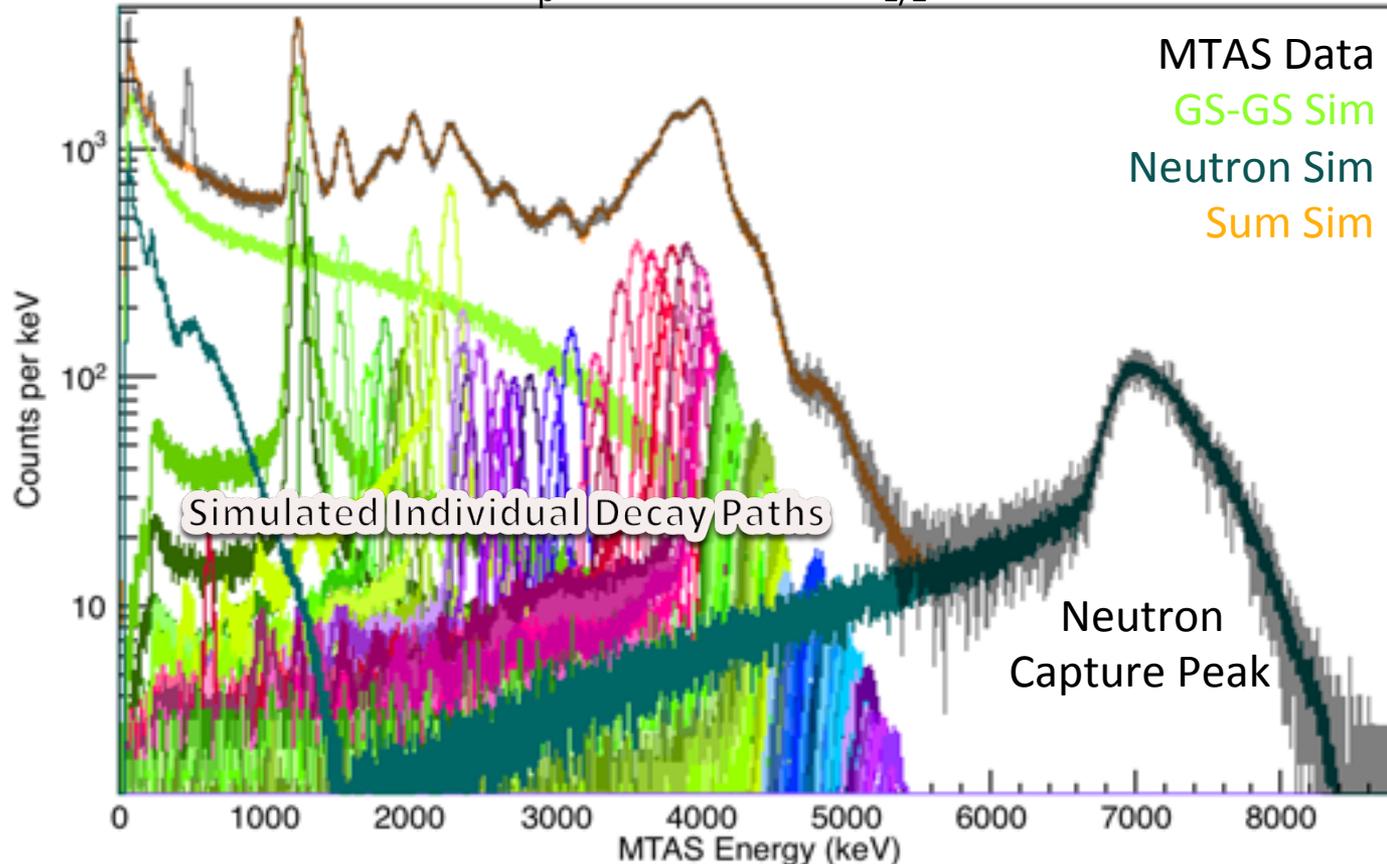
β -Feeding Intensity



Measuring All Three at Once

β s, γ s, and Neutrons: ^{137}I

^{137}I : $Q_{\beta} = 6027(8)$ keV $T_{1/2} = 24.5(2)$ s

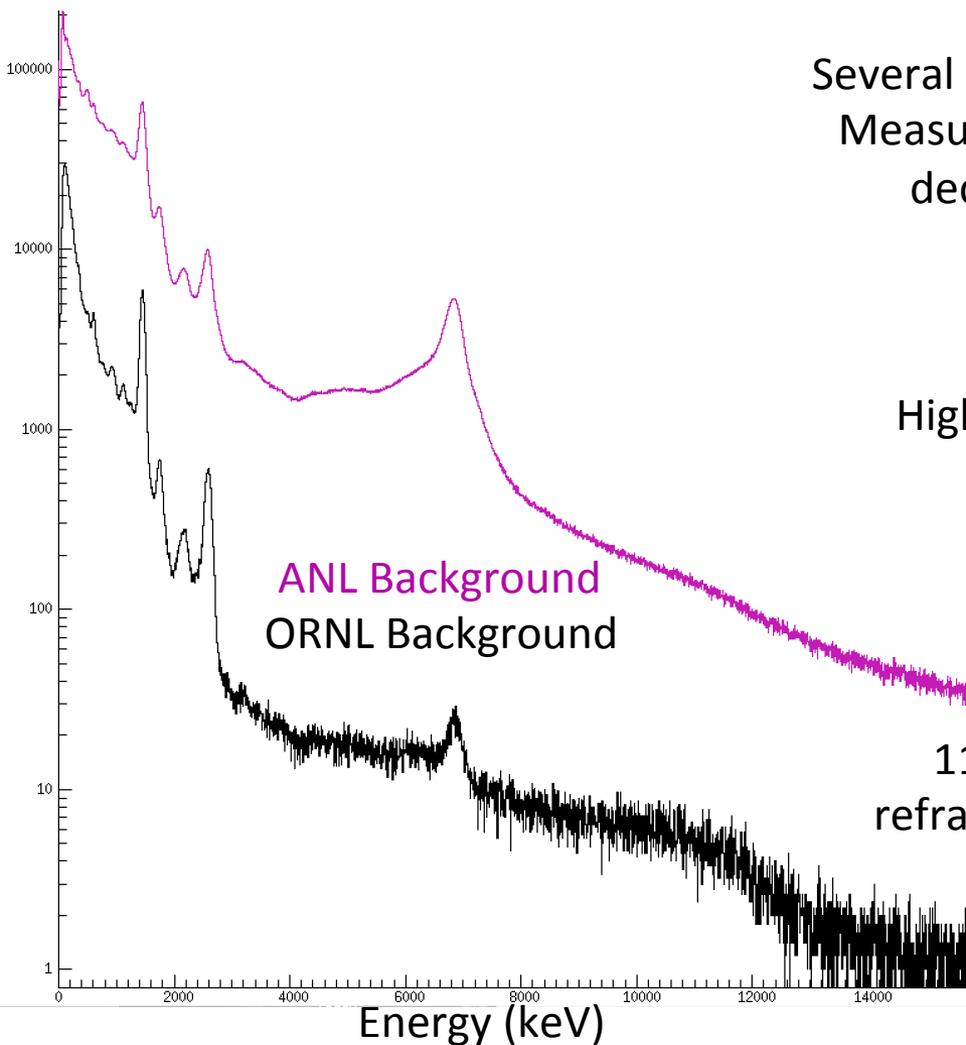


B.C. Rasco, *et al.*, PRC **95**, 054328 (2017)

Extracting all three components, β , γ , and neutrons from one decay
This decay is important for r process decay to stability

MTAS

Current and Previous Measurements



Several campaigns in 2012-2016 at the HRIBF at ORNL:
Measured over 70 neutron rich nuclei important for
decay heat and reactor antineutrino physics.
About a dozen have been evaluated.

2017-2018:

High precision source decay measurements of
 ^{88}Y , ^{54}Mn , ^{65}Zn , ^{40}K

Up next:

At Argonne: Arrived in March 2019
11 days of beam time for measurement of
refractory elements + some commissioning time

FRIB

Not equal measurement time

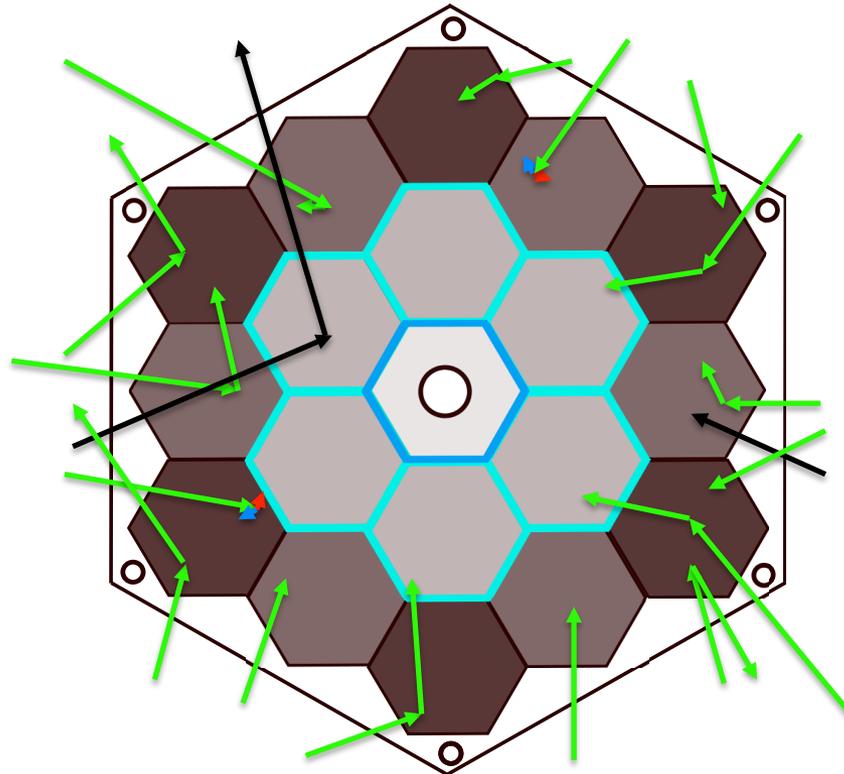
MTAS as an Active Veto

What are these measurements about?

2017-2018:

High precision source decay measurements of
 ^{88}Y , ^{54}Mn , ^{65}Zn , ^{40}K

MTAS Active Veto

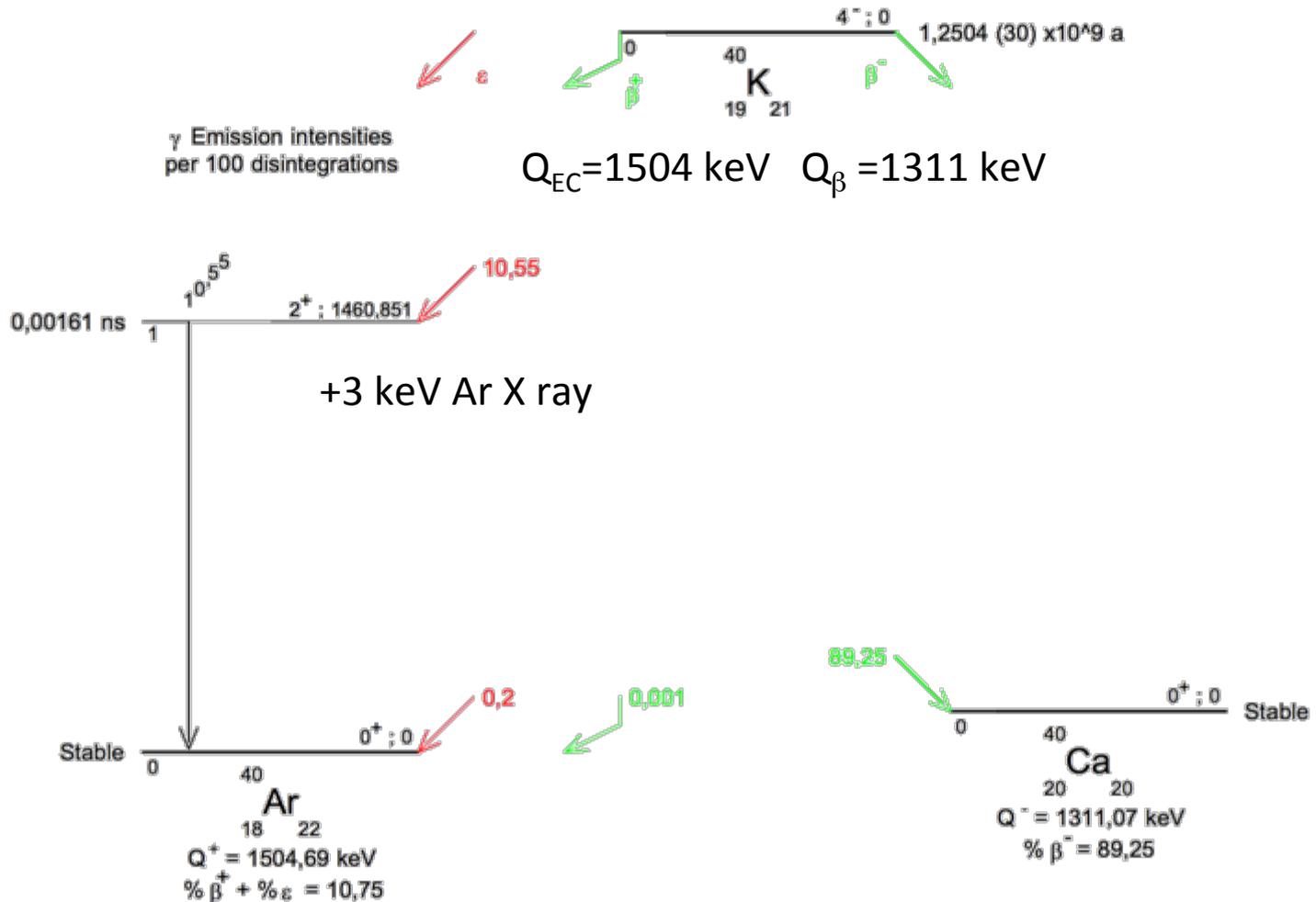


MTAS

Why ^{40}K Measurements?

Background for signal (3-6 keV bump) at DAMA, SABRE, COSINE-100,... and other high precision beyond the standard model physics measurements

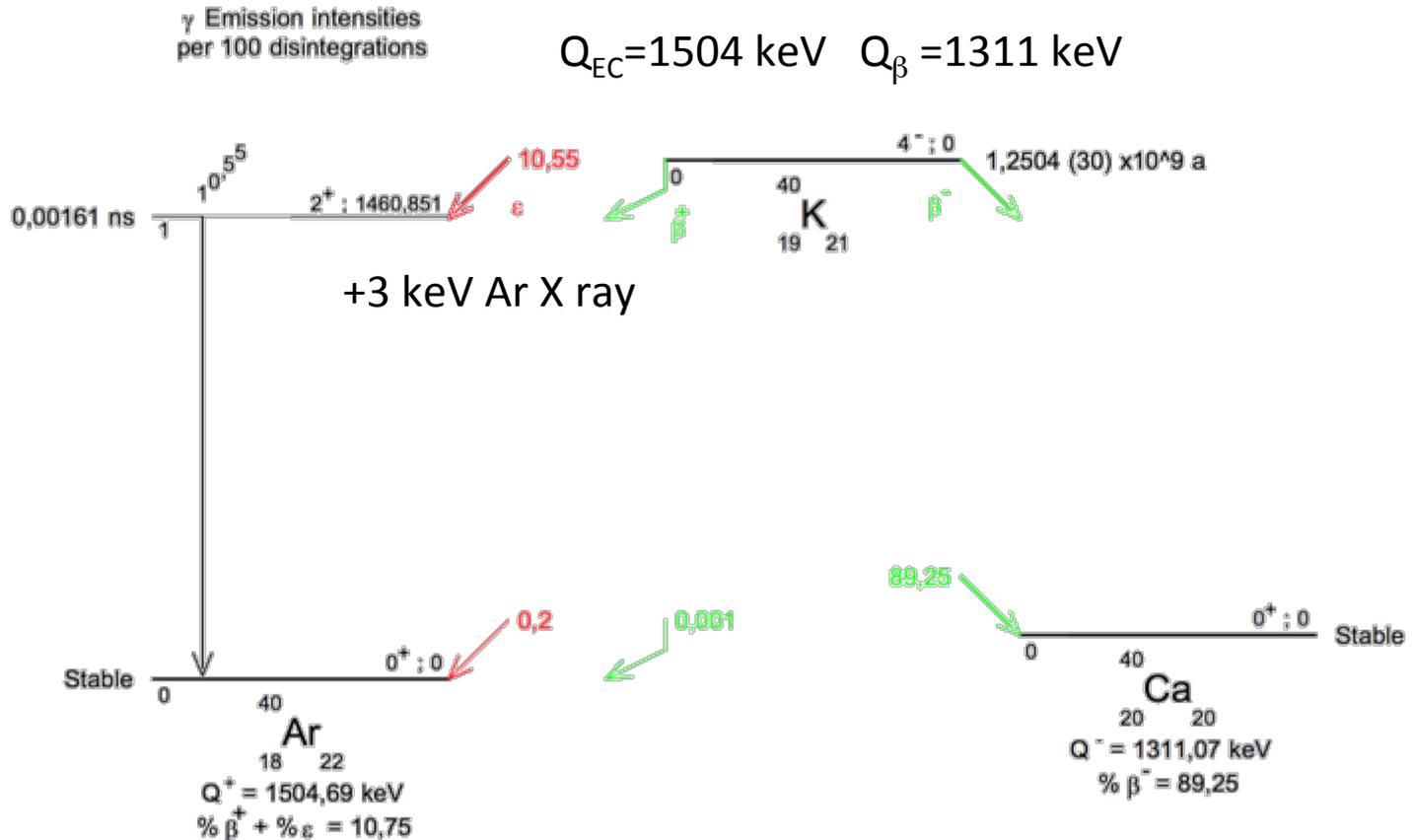
Improve precision of K-Ar Dating



MTAS

Why ^{40}K Measurements?

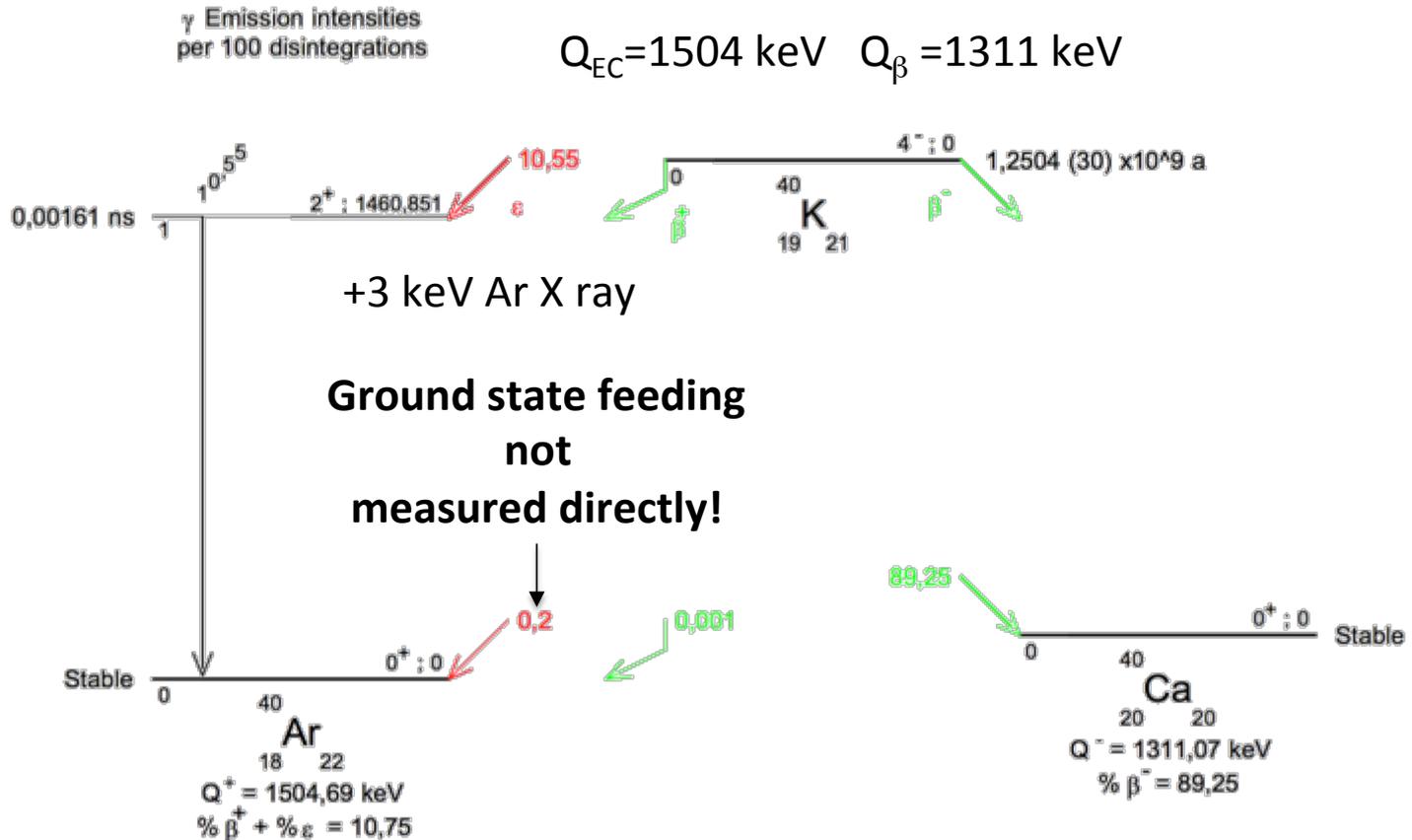
Background signal (3-6 keV bump) for DAMA, SABRE, COSINE-100,... and other high precision beyond the standard model physics measurements
 Improve precision of K-Ar Dating



MTAS

Why ^{40}K Measurements?

Background signal (3-6 keV bump) for DAMA, SABRE, COSINE-100,... and other high precision beyond the standard model physics measurements
 Improve precision of K-Ar Dating



^{40}K Decay to the Ground State

The different calculated branching ratios, $I_{(\beta^+ \epsilon)}$, of ^{40}K (EC)

LNE-LNHB/CEA : Table of Radionuclides Vol 5, ISBN 978-92-822-2234-8 (2010)

$$BR_{EC} = 0.2(1)\%$$

NNDC : Nuclear Data Sheets 140, 1 (2017)

$$BR_{EC} = 0.046(6)\%$$

Indirect Experimental Half-Life Value

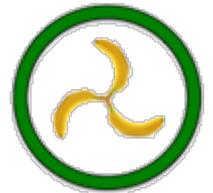
$$BR_{EC} = 0.8(8)\%$$

KDK Collaborator Value (J. Kostensalo and J. Suhonen, Jyväskylä)

$$BR_{EC} = 0.064(19)\%$$

Calculated one σ limits vary from $0.04\% < BR_{EC} < 0.3\%$
(~0-3% of electron capture measurement)

^{40}K Decay Experiment (KDK)

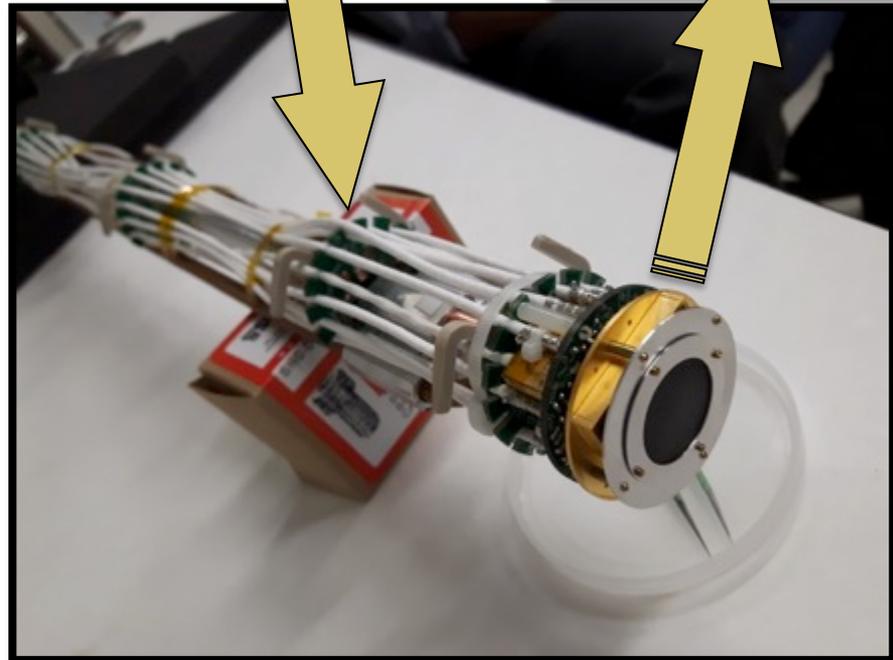
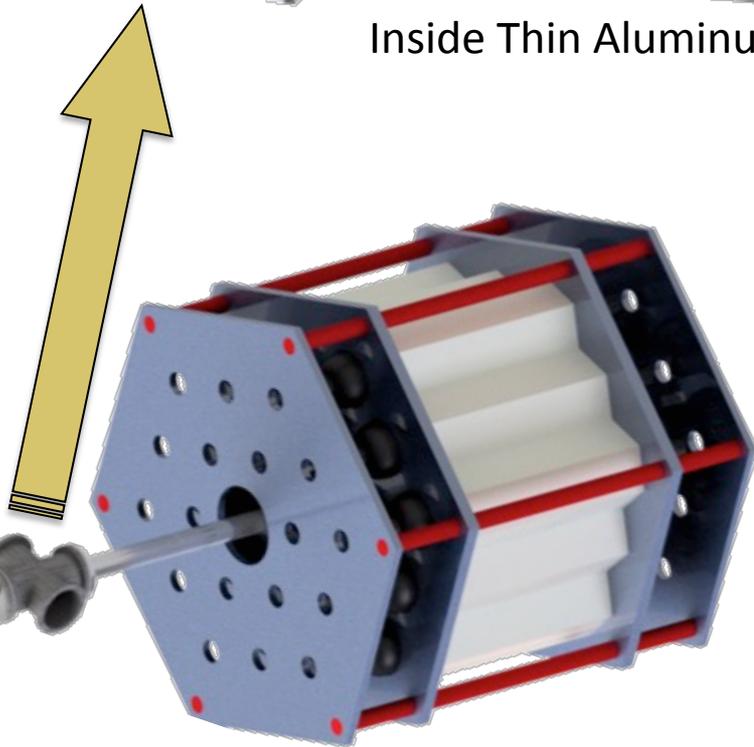
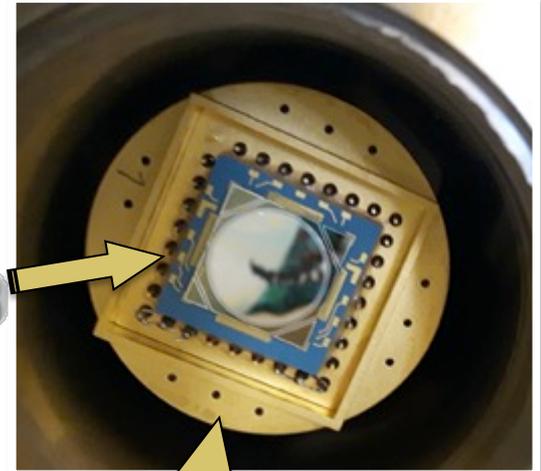


SDD X-Ray Detector

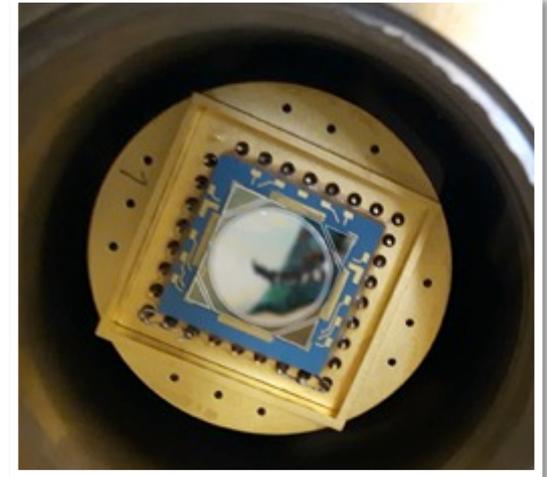
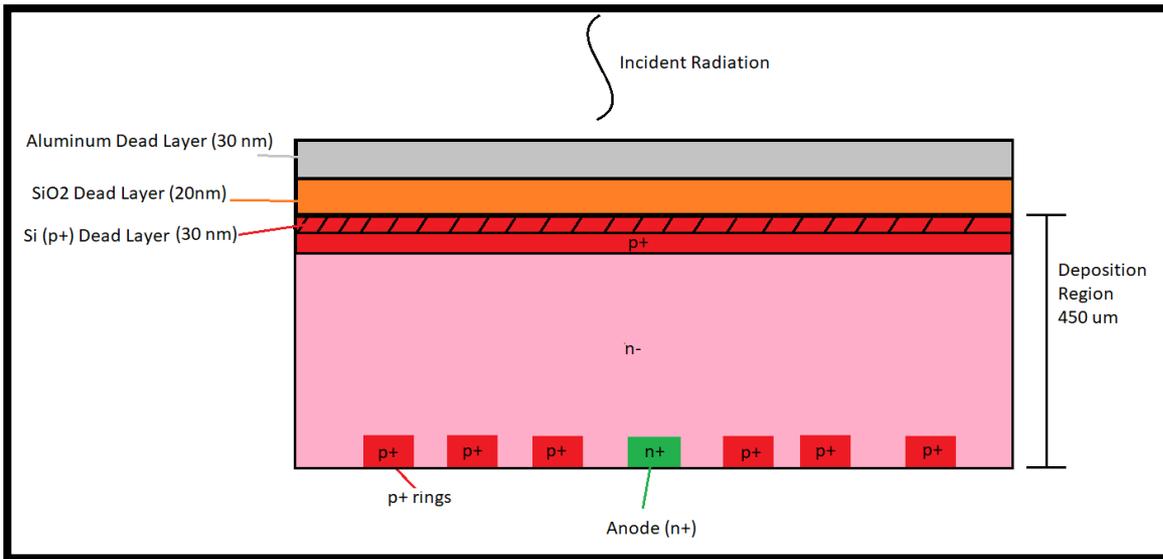
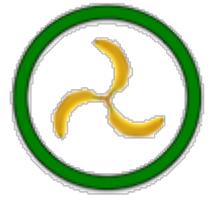


SDD Insert, Cooling, HV, and Readout

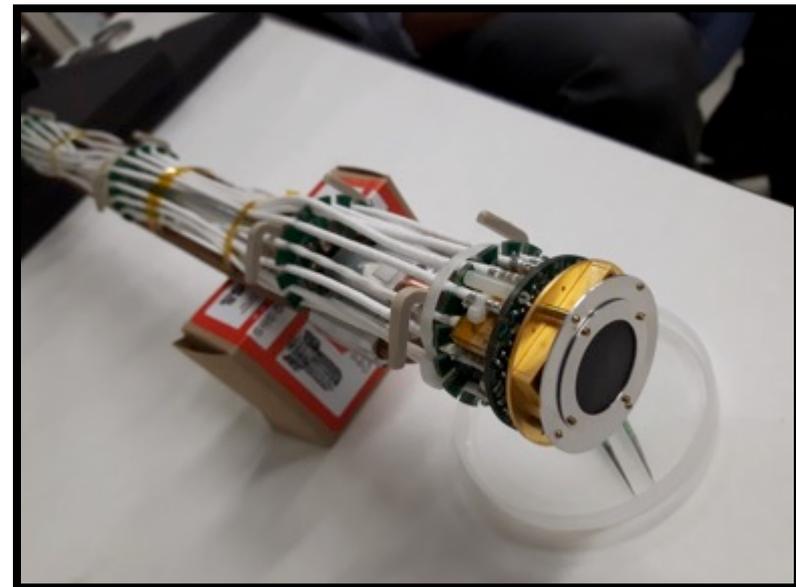
Inside Thin Aluminum Housing



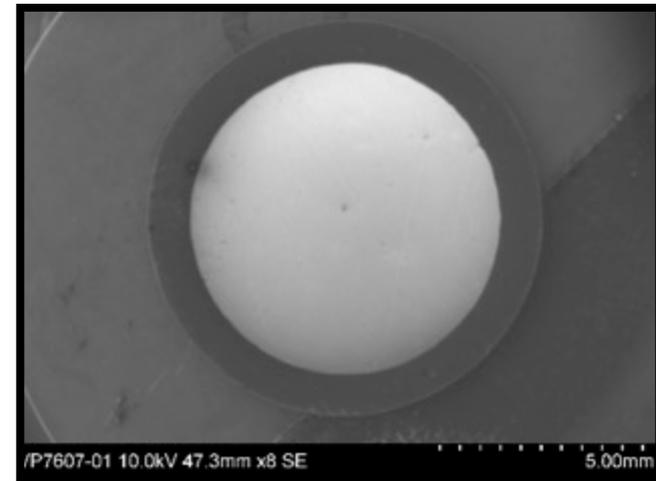
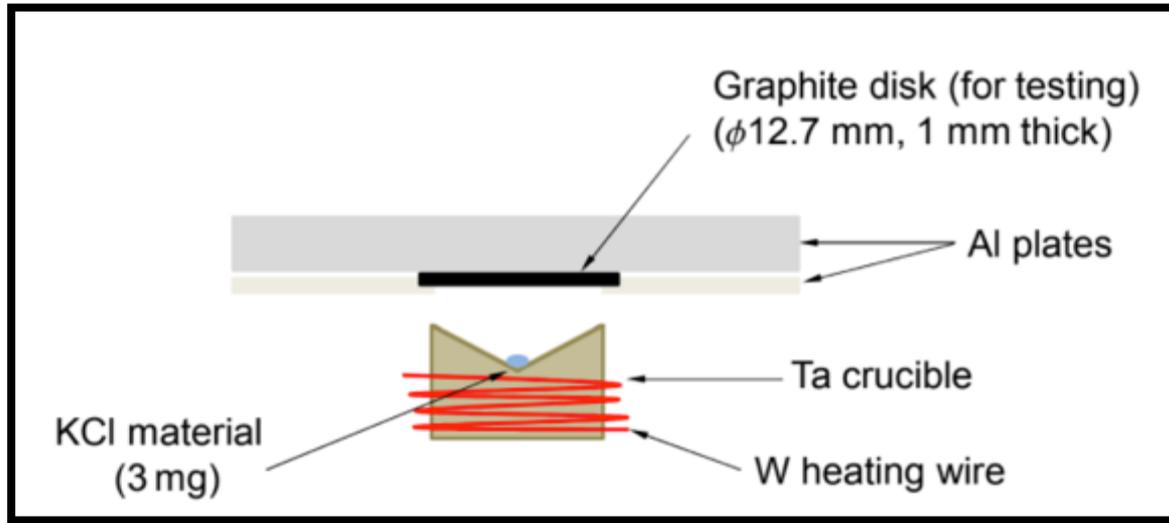
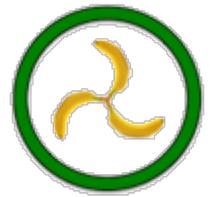
Silicon Drift Detector (SDD)



SDD Made at MPG, Munich
~15th iteration of design
They put SDDs on the Mars Rovers

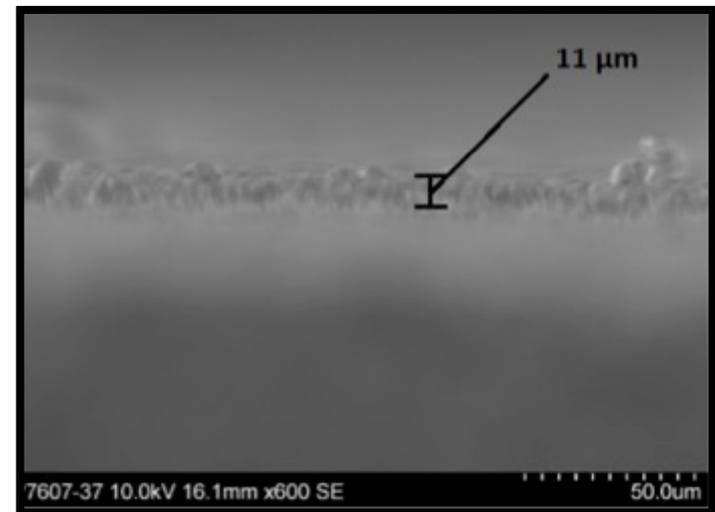


Making Enriched ^{40}K Source

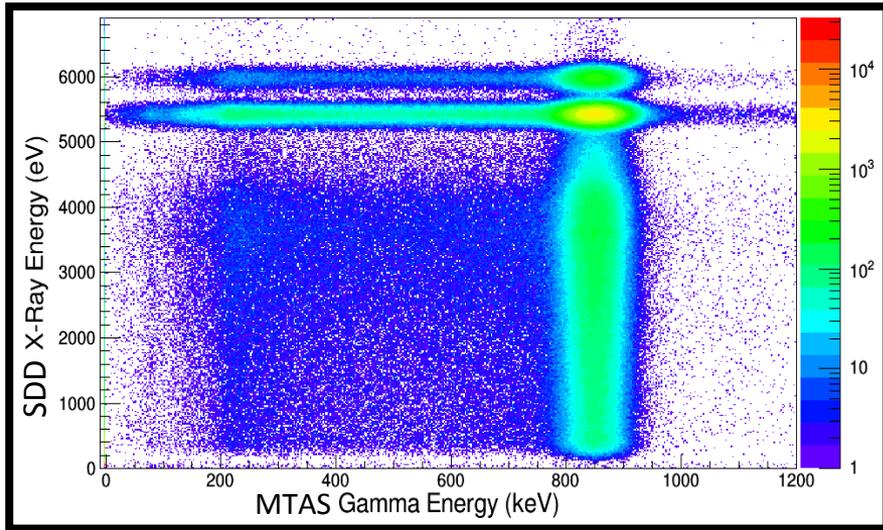
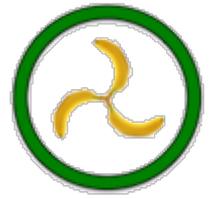


Enriched (16%) 3mg ^{40}KCl source evaporated
onto Carbon foil backing at ORNL
(~3 bananas equivalent)

Needs to be thin so 2.9 keV X rays
can escape from source



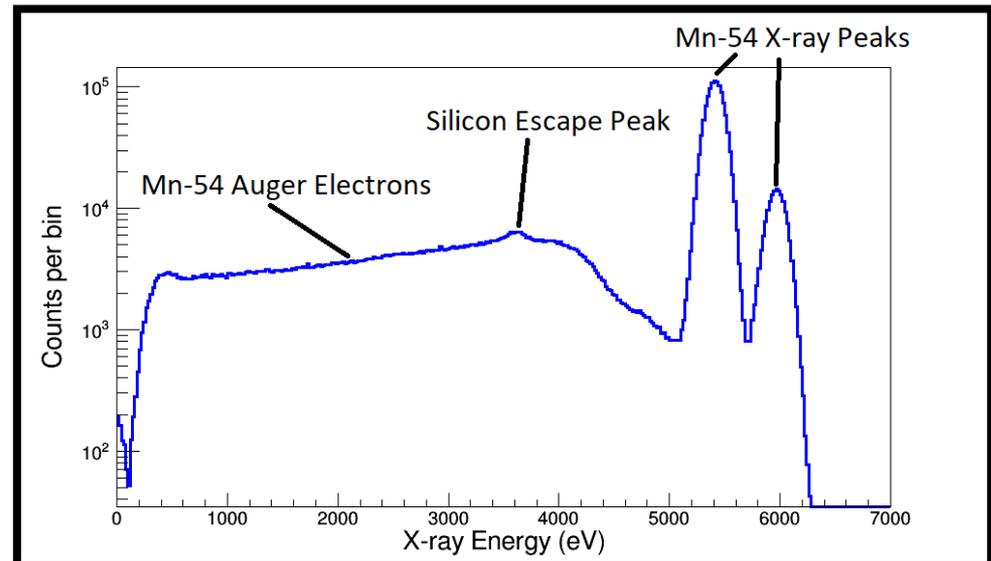
^{54}Mn in KDK: SDD and MTAS



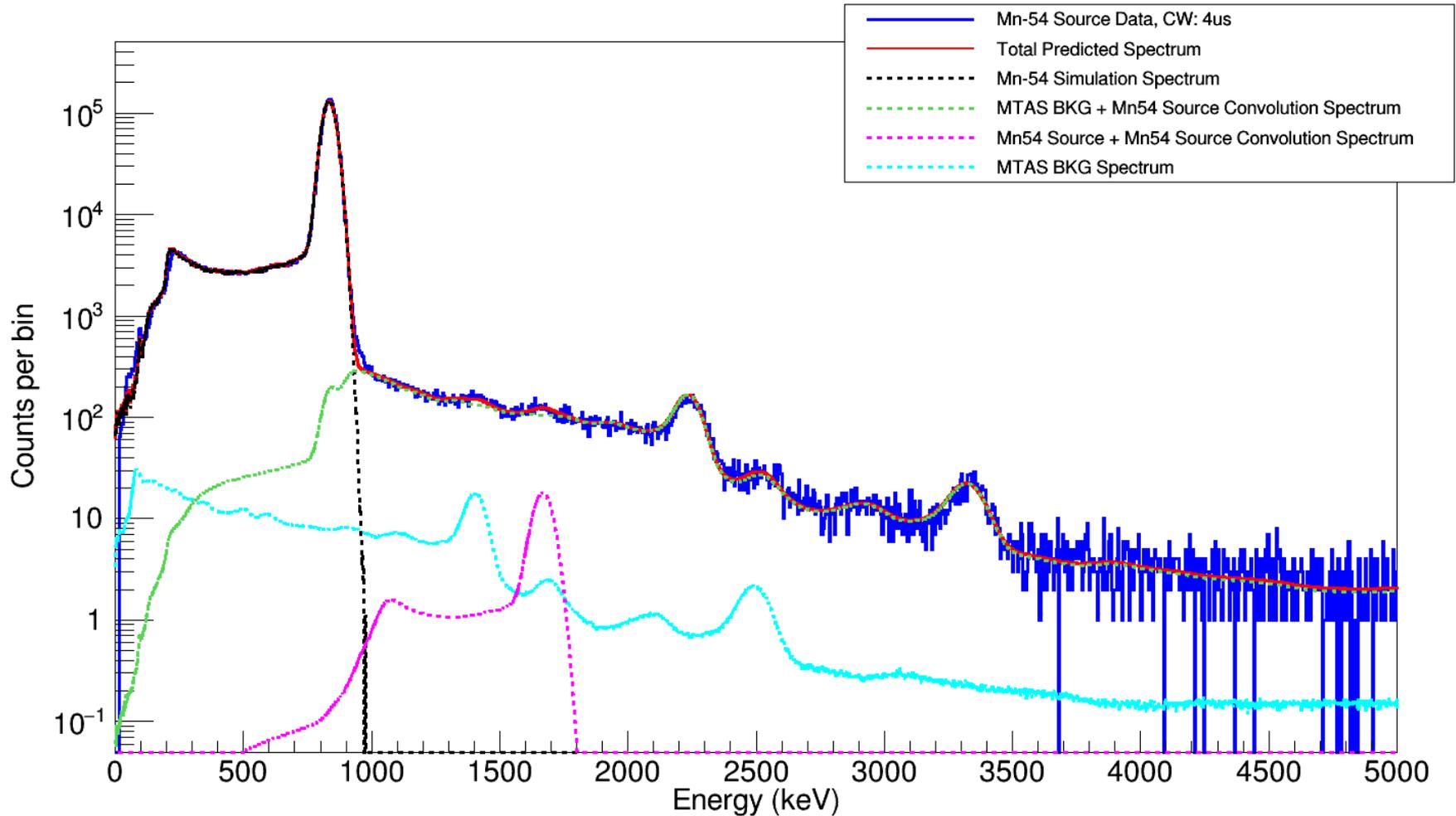
~50 nCi ^{54}Mn source
200 cps triggered in SDD

Gating on each ^{54}Mn X rays
gives consistent MTAS single
834 keV γ ray tagging
efficiency to
4 significant digits!

Finalizing dead-time corrections
before giving out efficiency

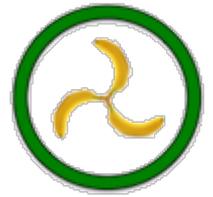


^{54}Mn with SDD Trigger in MTAS

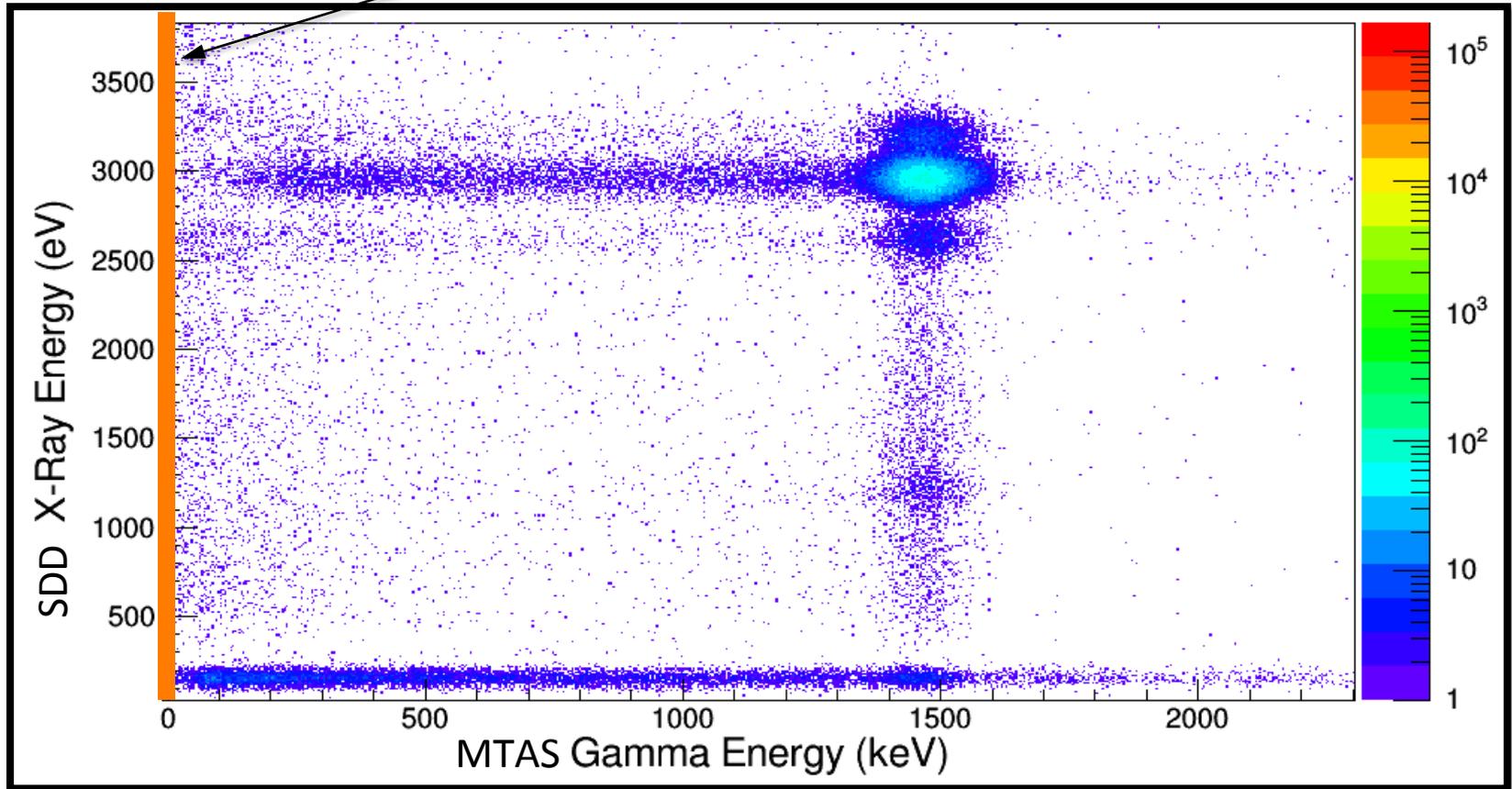


^{54}Mn spectrum in MTAS given a trigger in the SDD

^{40}K Preliminary MTAS

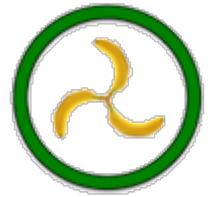


Blinded Data! Results soon!



33 days of data with about 1 triggered counts per minute

The KDK Collaboration



N.T. Brewer^[1], P. Di Stefano^[2], A. Fijalkowska^{[1][5][6]}, Z. Gai^[1], C. Goetz^[3],
R.Grzywacz^[3], J. Kostensalo^[7], P. Lechner^[8], Y. Liu^[1], E. Lukosi^[3], M. Mancuso^[9],
D. McKinnon^[3], C. Melcher^[3], J. Ninkovic^[8], F. Petricca^[9], B.C. Rasco^[1],
K.P. Rykaczewski^[1], P. Squillari^[2], D.W. Stracener^[1], M. Stukel^[2], J. Suhonen^[7],
M. Wolińska-Cichocka^{[1][4][6]}, Itay Yavin

- [1] Oak Ridge National Laboratory (ORNL), Tennessee, USA
- [2] Queen's University, Kingston, Ontario, CA
- [3] University of Tennessee at Knoxville, Knoxville, Tennessee
- [4] Heavy Ion Laboratory University of Warsaw, Warsaw, Poland
- [5] University of Warsaw, Warsaw, Poland
- [6] Joint Institute for Nuclear Physics and Applications (JINPA)
- [7] University of Jyväskylä, Jyväskylä, Finland
- [8] MPG Semiconductor Laboratory, Munich, Germany
- [9] Max Planck Institute for Physics, Munich, Germany

Technical and Electronic Support from M. Constable, F. Retiere (TRIUMF),
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arXiv: 1711.04004

Thank You
for your
Attention